Generating entangled spin states for quantum metrology by single-photon detection

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We propose and analyze a probabilistic but heralded scheme to generate pure, entangled, non-Gaussian states of collective spin in large atomic ensembles by means of single-photon detection. One photon announces the preparation of a Dicke state, while two or more photons announce Schrödinger cat states. The method produces pure states even for finite photon detection efficiency and weak atom-photon coupling. The entanglement generation can be made quasideterministic by means of repeated trial and feedback, enabling metrology beyond the standard quantum limit.

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I. INTRODUCTION

State-of-the-art atomic clocks and other atom interferometers are limited by quantum projection noise. For measurements on a system of N uncorrelated atoms in a coherent spin state (CSS), this projection noise sets a limit scaling as $1/\sqrt{N}$, referred to as the standard quantum limit (SQL). Entangled states can overcome this limit, potentially reaching the Heisenberg limit, where uncertainty scales as 1/N. Thus far, the potential for metrological gain has been demonstrated in atomic ensembles using squeezed spin states [1-9], which have enabled atomic clock operation surpassing the SQL [10,11]. In these experiments, the entanglement has been produced either by spin-dependent atom collisions [7-9] or by coupling an optical probe to the atomic ensemble [2-6]. The Greenberger-Horne-Zeilinger states [12] have also been shown to allow metrological gain [13] and have been produced [14,15] for collections of up to 14 ions [16] via Coulomb interactions.

In this paper, we describe a method to generate pure entangled states of collective spin in large atomic ensembles for measurements beyond the SQL. Photons transmitted through the ensemble experience a weak random Faraday rotation associated with the quantum noise of the atomic spin. A photon emerging with polarization orthogonal to its input polarization heralds the creation of an entangled state of collective atomic spin. The state thus produced is non-Gaussian in the sense that its phase-space representation on the Bloch sphere [17,18] is not described by a Gaussian distribution, unlike that of a CSS or even a moderately spin-squeezed state. Two or more orthogonally polarized photons emerging from the system herald increasingly more entangled "squeezed Schrödinger cat" states [19]. This method generates states of high purity even for weak atom-photon coupling and finite photon detection efficiency, which simply reduce the probability of entangledstate preparation. The heralded entanglement scheme can be made quasideterministic through repeated trial and feedback, enabling atom interferometry beyond the SQL.

Our approach has similarities to weak-measurement schemes [20] that use postselection to enable detection of small signals in the presence of technical noise [21-24]. It is closely

related to heralded schemes for the creation of Dicke states in atomic ensembles for quantum communication [25] which have been experimentally implemented for photon-pair and single-photon generation [26-28]. In our method, quantum noise of the atomic state produces a weak Faraday rotation of the polarization of a photon, whereby the phase of the atomic CSS becomes entangled with the photon polarization. The detection of a single photon of select polarization then prepares the atomic ensemble in a non-Gaussian entangled state that results from destructive interference between two weakly separated coherent states. This method can be implemented either in free space or in an optical cavity; the latter increases the polarization rotation and hence the entanglement rate. When the state preparation is fast compared to the atomic coherence time, as is the case in atomic clocks and many interferometers, then the present method can be made quasideterministic by repeated trial and feedback, enabling interferometry beyond the SQL. The metrological gain is 3 dB for just one detected photon and improves with additional detected photons. We note that related methods to generate squeezed Gaussian states by measuring the Faraday rotation of a light pulse containing a large number of photons have been proposed [29–31] and implemented [3]. The scheme proposed and investigated in Refs. [32,33] for Dicke-state preparation is similar to our scheme and should allow the same metrological gain.

II. STATE PREPARATION

Consider an ensemble of *N* three-level atoms. Two groundstate magnetic sublevels, $|\uparrow\rangle$, $|\downarrow\rangle$, correspond to a pseudospin $s_i = \frac{1}{2}$. The collective state of the ensemble can be described by a total spin $\mathbf{S} = \sum \mathbf{s_i}$ that is the sum of individual spins $\mathbf{s_i}$. Two degenerate, oppositely circularly polarized modes of an optical cavity couple $|\uparrow\rangle$ and $|\downarrow\rangle$ to an excited state $|e\rangle$ (Fig. 1). A magnetic field applied along the quantization axis \hat{z} lifts the degeneracy between the ground states by an amount $\hbar\delta$ such that the two-photon Raman coupling between $|\uparrow\rangle$ and $|\downarrow\rangle$ is negligible. We assume that the two transitions have equal coupling strength and that all atoms are equally coupled to the light, with single-photon Rabi frequency 2g. When the light-atom detuning Δ is much larger than the excited-state



FIG. 1. (Color online) Scheme for the heralded generation of nonclassical states. (a) Atoms with two spin states $|\uparrow\rangle$ and $|\downarrow\rangle$ are coupled to an electronic excited state $|e\rangle$ via two degenerate circularly polarized modes. (b) Incident vertically polarized photons experience weak Faraday rotation as they traverse the ensemble. The detection of a horizontally polarized transmitted photon heralds the generation of a non-Gaussian entangled state of collective atomic spin. A cavity enhances the Faraday rotation and the state preparation probability.

width Γ , we can adiabatically eliminate the excited state $|e\rangle$. Ignoring photon emission into free space for now, the interaction Hamiltonian for the atom-photon system is written as [29]

$$\frac{H}{\hbar} = \left(\frac{2g^2}{\Delta}\right) J_z S_z. \tag{1}$$

J is the Stokes vector of light and obeys the commutation relation of angular momenta $[J_i, J_j] = i\epsilon_{ijk}J_k$. In particular, $J_z = \frac{1}{2}(a^{\dagger}_+a_+ - a^{\dagger}_-a_-)$ where a_{\pm} are the annihilation operators of σ^{\pm} light. The atoms are prepared initially in the CSS $|x\rangle$ along \hat{x} , satisfying $S_x|x\rangle = S|x\rangle$ where S = N/2. Consider a vertically polarized incident photon described by the state $|v\rangle = (|\sigma^+\rangle + |\sigma^-\rangle)/\sqrt{2}$. While the photon is inside the cavity, the atom-photon system evolves as $e^{-iHt}|x\rangle|v\rangle$, which after the photon has been transmitted through the cavity results in the state [34]

$$\begin{aligned} |\psi_t\rangle &= \frac{1}{\sqrt{2}} \sum_{m=-S}^{S} c_m |m\rangle (e^{-im\phi} |\sigma^+\rangle + e^{im\phi} |\sigma^-\rangle) \\ &= \frac{1}{\sqrt{2}} (|\sigma^+\rangle |\phi\rangle + |\sigma^-\rangle |-\phi\rangle). \end{aligned}$$
(2)

Here $\phi = \eta \Gamma/2\Delta$ is an accumulated phase, expressed in terms of the cavity linewidth κ and the single-atom cooperativity $\eta = 4g^2/\kappa\Gamma$ [34,35], and the atomic state is written in terms of S_z eigenstates $|m\rangle$ and binomial coefficients $c_m = 2^{-S}\{(2S)!/[(S+m)!(S-m)!]\}^{1/2}$. Here $|\pm\phi\rangle$ designates the CSS in the equatorial plane rotated by an angle $\pm\phi$ about \hat{z} away from \hat{x} . In the following we restrict the analysis to weak atom-cavity coupling, $\eta \ll 1$, and the dispersive limit of low photon absorption, requiring [35] $2S\eta(\Gamma/2\Delta)^2 \ll 1$. This implies that the angle ϕ is much smaller than the CSS angular width $\phi_{css} = 1/\sqrt{2S}$.

The first line of Eq. (2) is readily interpreted as the phase $\pm m\phi$ being imprinted onto the σ^{\pm} polarizations of the light field due to the refractive index of the atoms in the states $|\uparrow\rangle$, $|\downarrow\rangle$ with population difference 2m. A value $m \neq 0$, i.e., a deviation of S_z from its mean value $\langle S_z \rangle = 0$ due to quantum noise in the atomic state, thus results in a polarization rotation of the photon. The detection of a horizontally polarized



FIG. 2. (Color online) Normalized Wigner quasiprobability distribution $W(\theta, \phi)/\sqrt{2\pi S}$ (left) and probability distributions of angular-momentum eigenvalues (right, solid line for S_z , dashed line for S_y) calculated for N = 50 atoms for [(a), (b)] the input CSS, [(c), (d)] n = 1 detected |h⟩ photon, [(e), (f)] n = 2 detected |h⟩ photons, and [(g), (h)] n = 5 detected |h⟩ photons. Wigner functions [(e), (g)] indicate the production of a Schrödinger cat state for $n \ge 2$ detected |h⟩ photons. The distributions of S_y eigenvalues in panels (d), (f), and (h) consist of several peaks, narrower than the CSS width, enabling measurements surpassing the SQL.

photon $|h\rangle$ requires $S_z \neq 0$ and biases the system towards states with larger $|S_z|$, creating a collective spin state whose quasiprobability distribution on the Bloch sphere shows a hole in the center (see Fig. 2).

From a complementary viewpoint, σ^+ and σ^- photons shift the phase of the atomic CSS in opposite directions by an amount $\pm \phi$. Even though $\phi \ll \phi_{css}$, the detection of a horizontally polarized photon $|h\rangle = (|\sigma^+\rangle - |\sigma^-\rangle)/\sqrt{2}$ corresponds [according to the second line in Eq. (2)] to the destructive interference $|\phi\rangle - |-\phi\rangle$ of two weakly separated coherent spin states, which generates the hole in the center of the state.

The atomic state after detection of one photon in $|h\rangle$ can be expressed as $|\psi_1\rangle = \sum_m c_m \sin(m\phi)|m\rangle$, where the unimportant normalization factor has been omitted, and for $\phi \ll \phi_{css}$ approximated as $|\psi_1\rangle = \sum_m mc_m |m\rangle$. $|\psi_1\rangle$ is the first Dicke state along \hat{x} , which satisfies $S_x |\psi_1\rangle = (S-1)|\psi_1\rangle$.

We now consider the effect of a multiple-photon input state. We begin by considering an input Fock state and later discuss the effects due to the more easily realizable situation of an input coherent state. For an input photon Fock state of n_0 photons, with *n* photons exiting the system in $|h\rangle$ and $n_0 - n$ photons exiting in the original polarization $|v\rangle$, the atomic state is given by

$$|\psi_n\rangle = \sum_m c_m \sin^n(m\phi) \cos^{n_0 - n}(m\phi) |m\rangle.$$
(3)

In the dispersive limit and for small η , $\cos(m\phi) \approx 1$ for $m \lesssim \sqrt{S/2}$, and the state $|\psi_n\rangle$ for $n \ge 2$ corresponds to a "squeezed cat" state [19], a superposition of two Gaussian states squeezed by a factor of 2 and separated on the Bloch sphere by an angle $\Delta\theta = 2\sqrt{n/S} = \sqrt{8n}\phi_{css}$ (see Fig. 2). Remarkably, as both the separation angle $\Delta\theta$ and the CSS angular width ϕ_{css} scale as $1/\sqrt{S}$, this allows the production of states separated by an angle greater than the CSS width for just a few detected photons, regardless of the atom number used. We emphasize that for $\phi \ll \phi_{css}$ the states $|\psi_n\rangle$ are independent of ϕ , which affects only the likelihood of producing the state.

III. METROLOGICAL GAIN

The entangled states $|\psi_n\rangle$ display peaks in the angularmomentum distributions along both S_z and S_y that are narrower than the CSS width. In particular, expressed in terms of S_y eigenstates $|m\rangle_y$, the state $|\psi_n\rangle$ is to lowest order independent of n_0 and is well approximated by

$$|\psi_n\rangle = \begin{cases} \sum_m A_{n,S} e^{\frac{-m^2}{4S}} \sin(m\sqrt{n/S}) |m\rangle_y & n \text{ odd,} \\ \sum_m A_{n,S} e^{\frac{-m^2}{4S}} \cos(m\sqrt{n/S}) |m\rangle_y & n \text{ even,} \end{cases}$$
(4)

where $A_{n,S} = (\pi S/2)^{-1/4}(1 - e^{-2n})^{-1/2}$ is a normalization constant. Figure 2 shows the S_z and S_y probability distributions of the state $|\psi_n\rangle$. We also plot the corresponding Wigner function $W(\theta, \phi)$ for the spin state, given by [36]

$$W(\theta,\phi) = \sum_{k=0}^{2S} \sum_{q=-k}^{k} \rho_{kq} Y_{kq}(\theta,\phi), \qquad (5)$$

where the terms ρ_{kq} represent density matrix elements in the spherical harmonic basis and the $Y_{kq}(\theta,\phi)$ are the spherical harmonics. We note that the probability distributions of angular momentum eigenvalues are given by the marginals of the Wigner function. Figures 2(a) and 2(b) show the input CSS, while Figs. 2(c) and 2(d) show the state produced by the conditional detection of one photon in $|h\rangle$. Higher-order states produced by the conditional detection of more than one $|h\rangle$ photon are shown in Figs. 2(e)–2(h).

The narrower features along S_y enable improved phase readout in a Ramsey measurement compared to the CSS: After initial-state preparation, a period τ of free evolution is followed by rotation about the average direction of the spin vector, which can be chosen as \hat{x} , thereby mapping the multipeaked S_y distribution onto the S_z axis. The value of $\langle S_z \rangle$, and hence the accumulated interferometer phase, is found by fitting the measured distribution of S_z values to the *a priori* distribution given by Eq. (4). (We assume that decoherence in the interferometer leads to phase fluctuations much less than the width of the peaks.) To see that this



FIG. 3. Black squares show the measurement variance of the states $|\psi_n\rangle$, normalized to the CSS variance, $(\Delta S_z^{(w)})^2/(S/2)$, in dB, as a function of number of detected $|h\rangle$ photons *n*, indicating substantial metrological gain for *n* of a few. The dashed curve shows that for larger *n* the normalized variance asymptotically approaches the value $(\Delta S_z^{(w)})^2/(S/2) = 0.64/n$ (see text). Due to the finite atom number, N = 100, used for the calculations, points in the figure show slight deviations from the asymptotic behavior expected for large *N*.

procedure gives lower quantum uncertainty in the estimation of $\langle S_{\tau} \rangle$, consider M measurement points, of which a known fraction f_i falls under a particular peak i, μ_i is the mean value of S_z associated with that peak, and σ_0 is the width of each peak. The weighted average, given by $S_z^{(w)} = \sum_i \mu_i f_i$, has variance given by $(\Delta S_z^{(w)})^2 = \sigma_0^2 / M$, the same as for M measurements conducted on a single peak of width σ_0 . Thus, a probability distribution composed of multiple narrow peaks allows the same reduction in measurement uncertainty as one containing a single peak of equal (reduced) width. This allows the entangled states $|\psi_n\rangle$ to produce substantial metrological gain. In particular, the first excited Dicke state, produced by a single detected photon, results in measurement variance 3.4 dB below the SQL. This metrological gain is confirmed by calculations of the classical Fisher information in the S_z distributions [37], which show enhancement beyond the SQL in agreement with the values obtained by the measurement protocol we have described.

Figure 3 shows the measurement variance (normalized to the CSS variance) as a function of detected photon number n, assuming large atom number N. The normalized variance asymptotically approaches the value 0.64/n (dashed curve in Fig. 3). This represents the squared ratio of the width of one peak of the function $\cos^2(m\sqrt{n/S})$, assuming n large, to the CSS width $\sqrt{S/2}$. While the probability to produce high-nstates decreases exponentially, Fig. 3 indicates that substantial metrological gain is obtained even for n of a few. (Similar calculations for higher-order excited Dicke states [32] indicate that the nth Dicke state results in the same metrological gain as the n-photon state produced by our scheme.) We also note that metrological gain can be used to prove that the atomic ensemble is entangled [38,39].

IV. RATE OF ENTANGLED-STATE GENERATION

To calculate the heralded generation rate of these entangled states, we note that the probability of converting one incident $|v\rangle$ photon into an $|h\rangle$ photon and detecting it is easily

calculated from the mean square of the polarization rotation angle $\langle \beta^2 \rangle = S\phi^2/2$ and is given by $p = qS\phi^2/2 \ll 1$, where $q \leq 1$ is the photon detection efficiency. The probability of the incident photon being scattered into free space by the atomic ensemble is $p_{sc} = 2S\eta(\Gamma/2\Delta)^2 = 2S\phi^2/\eta$ [35]. Therefore the success probability is simply related to the free-space scattering probability via $p = q\eta p_{sc}/4$. A cavity increases the single-atom resonant optical depth 2η [35] and hence greatly improves the generation efficiency for a given p_{sc} .

In realistic implementations, the input photons will typically be in a coherent state with mean photon number n_0 rather than a Fock state. This means that an unknown number of photons will exit the system in the original polarization $|v\rangle$. Whether detected or not, these $|v\rangle$ photons have minimal impact on the atomic state: They multiply the coefficients c_m of the state by $\cos(m\phi) \approx 1$ as shown in Eq. (3). The series of cosine factors does not significantly degrade the atomic state until $Nn_0\phi^2$ approaches unity. Since the probability p_1 to detect one outgoing $|h\rangle$ photon out of n_0 incident $|v\rangle$ photons is given by $p_1 = q Nn_0\phi^2/4$, this results in the requirement $p_1 \ll q$. Under this condition, an input coherent state of the light will have the same effect as an input Fock state of the light.

The produced state can also be degraded by undetected $|h\rangle$ photons. When a coherent state is used for the input, the probability to detect exactly *n* photons in $|h\rangle$ is given by the weighted sum over values of S_7 of $P(\bar{n}(S_7), n)$, where $P(\bar{n}(S_z), n)$ is the chance to find exactly *n* output photons given a Poisson distribution with mean value $\bar{n}(S_z) = q n_0 S_z^2 \phi^2$. For $p_1 \ll 1$, the overall probability for exactly *n* photons to exit the system in |h) is $(p_1/2)^n (2n)!/(n!)^2 q^n$, and the probability to detect them all is $p_n = (p_1/2)^n (2n)!/(n!)^2$. The probability that n + 1 photons exit in $|h\rangle$ of which n are detected is then $p_{n+1}q^n(1-q)(n+1)/q^{n+1} = p_1p_n(2n+1)(1-q)/q$. Such "false positive" states, corresponding to an additional undetected $|h\rangle$ photon, produce an atomic state different from the heralded state, substantially reducing the signal-to-noise ratio. Under the requirement $p_1 \ll q/(2n+1)$, the probability for such "false-positive" states is smaller than that of the heralded state by a factor $p_1(2n+1)(1-q)/q \ll 1$.

Note that, in order to maintain the coherence of the atomic spin state, the photon number scattered into free space must remain substantially smaller than the atom number, $n_0 p_{sc} \ll N$. For optically dense ensembles in free space, $2N\eta > 1$, this condition is automatically met by $p_1 \ll 1$, and the method proposed here can also be directly applied to dense ensembles in free space.

Given a coherent input state with mean photon number n_0 , the probability per trial to detect *n* photons is $(p_1/2)^n (2n)!/(n!)^2$, which requires on average only a small number $n_{sc} = 4p_1/q\eta$ of photons to be scattered into free space. While the success probability decreases exponentially with *n*, even states corresponding to *n* of a few display significant nonclassicality. For instance, with a realistic q =0.5 detection efficiency and choosing $np_1 = 0.2q$, the creation of entangled states corresponding to n = 1,2,3 requires on average 10, 300, and 1×10^4 trials, respectively, for any number of atoms. The corresponding improvements over the SQL are 3.4, 6.0, and 7.4 dB, respectively.

Although the states are generated only probabilistically, due to the heralding, preparation attempts may be repeated until success. For Ramsey measurements, the free precession time τ is typically much longer than the state preparation time. Under these conditions, even if entangled-state preparation requires many attempts, the total preparation time can remain small compared to τ and there is no significant reduction of measurement duty cycle. As one example, state preparation in the Sr optical lattice clock is limited by the 20 μ s decay time of the optical pumping transition, suggesting that state preparation, consisting of optical pumping to $|\downarrow\rangle$, preparation of a coherent state $|x\rangle$ with a $\pi/2$ pulse, and illumination by the probe pulse, could be performed in as little as 200 μ s, while the available measurement time for that system is ~ 1 s and is currently laser-limited [40]. Thus, up to $\sim 10^4$ state preparation attempts may be made without compromising available measurement time; a transition with faster excitedstate decay time may allow even more attempts.

V. CONCLUSIONS

In conclusion, we have proposed a technique where a single photon can create an entangled spin state of a very large ensemble of atoms. This can be achieved even in the limit of weak coupling between a photon and an atom and finite photon detection efficiency. The use of these states for interferometry below the SQL requires state readout capabilities well below the CSS width, as have been recently demonstrated [41]. While spin squeezing [1–9] can generate larger metrological gain than the method proposed here, all such experiments to date have produced antisqueezing that is much larger than the squeezing, resulting in mixed spin states. In contrast, the method proposed here generates nearly pure states and can be used to create "Schrödinger cat" states that are of fundamental interest.

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